## **CS5670: Computer Vision**

Reflectance & Photometric stereo







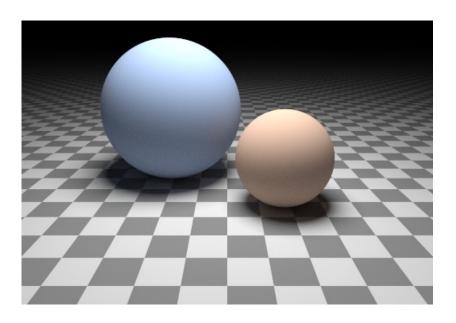
## Reading

• Szeliski 2<sup>nd</sup> Edition: Chapter 2.2 & 13.1

#### Roadmap for the rest of the course

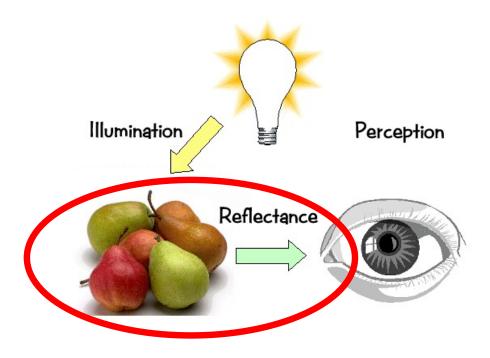
- The next three lectures will finish up geometry and image formation
  - Next up (after Spring Break): deep learning, image recognition, neural radiance fields, image generation models
- Coming up
  - Reflectance and Photometric Stereo (today)
  - Two-view geometry
  - Multi-view geometry

#### Can we determine shape from lighting?



- Are these spheres?
  - Or just flat discs painted with varying albedo?

#### Reflectance



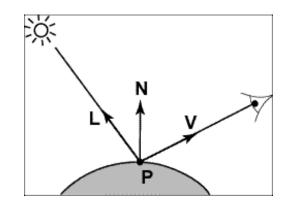
### **Light sources**

- Basic types
  - point source
  - directional source
    - a point source that is infinitely far away
  - area source
    - a union of point sources
- What happens when light hits an object?

#### **Modeling Image Formation**

#### We need to reason about:

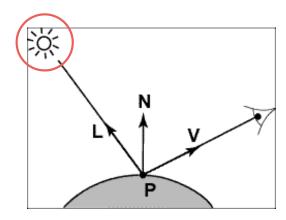
- How light interacts with the scene
- How a pixel value is related to light energy in the world



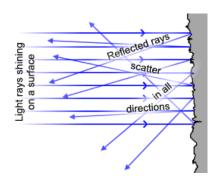
Track a "ray" of light all the way from light source to the sensor

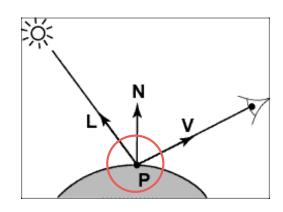
## **Directional Lighting**

- Key property: all rays are parallel
- Equivalent to an infinitely distant point source



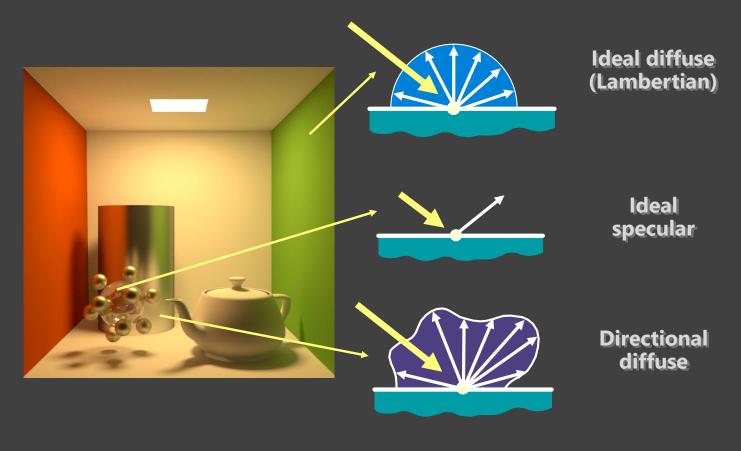
#### **Lambertian Reflectance**



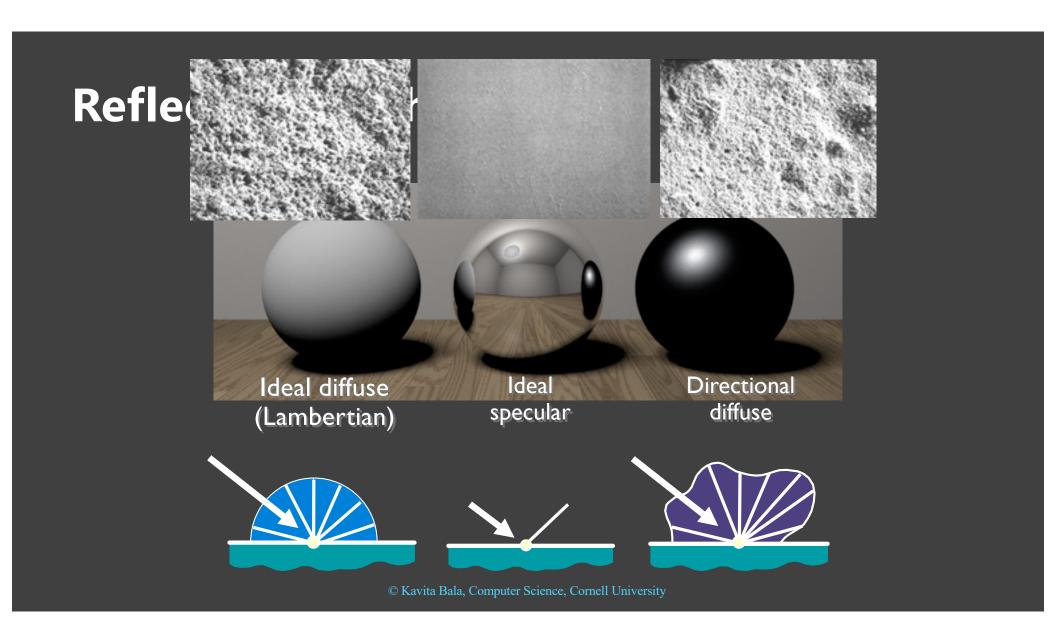


$$I = N \cdot L$$
 Image Surface Light intensity normal direction Image intensity  $\propto \cos(\text{angle between N and L})$ 

#### **Materials - Three Forms**

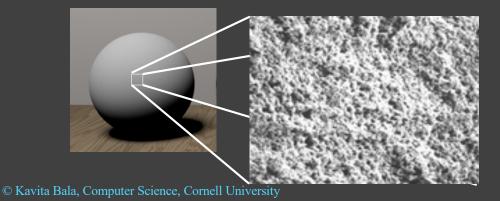


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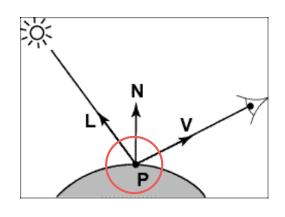


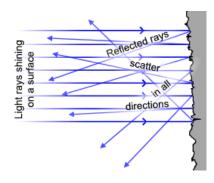
#### **Ideal Diffuse Reflection**

- Characteristic of multiple scattering materials
- An idealization but reasonable for matte surfaces



#### **Lambertian Reflectance**

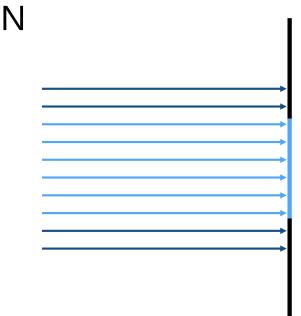




- 1. Reflected energy is proportional to cosine of angle between L and N (incoming)
- 2. Measured intensity is viewpoint-independent (outgoing)

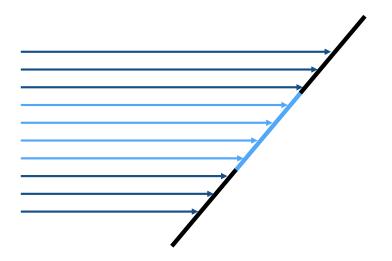
## Lambertian Reflectance: Incoming

 Reflected energy is proportional to cosine of angle between L and N



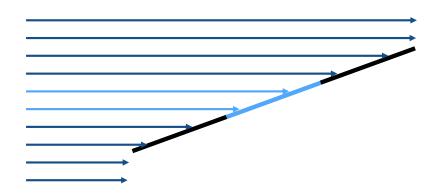
## Lambertian Reflectance: Incoming

 Reflected energy is proportional to cosine of angle between L and N



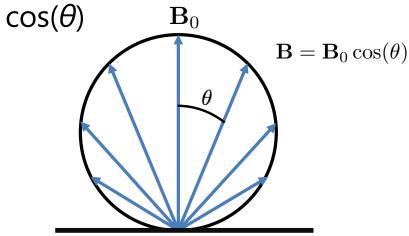
## Lambertian Reflectance: Incoming

 Reflected energy is proportional to cosine of angle between L and N



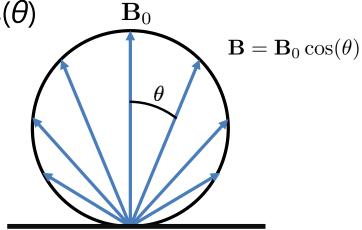
Light hitting surface is proportional to the **cosine** 

• Number of photons reflected to a given angle  $\theta$  is proportional to



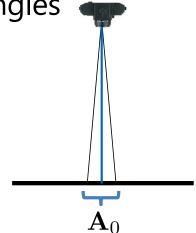
Lambert's cosine law: $\mathbf{B} = \mathbf{B}_0 \cos(\theta)$ 

• Number of photons reflected to a given angle  $\theta$  is proportional to  $\cos(\theta)$   $\mathbf{B}_0$ 

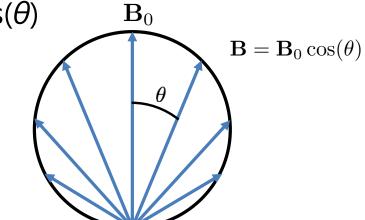


Lambert's cosine law:  $\mathbf{B} = \mathbf{B}_0 \cos(\theta)$ 

 But appearance is the same from every angle due to larger pixel footprint at larger angles

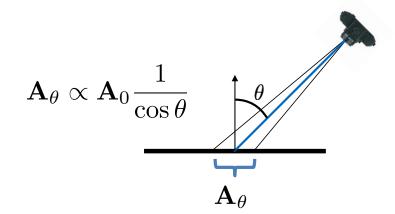


• Number of photons reflected to a given angle  $\theta$  is proportional to  $\cos(\theta)$   $\mathbf{B}_0$ 

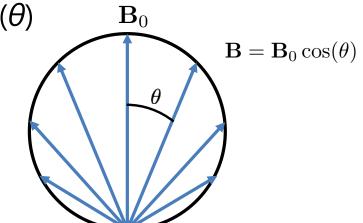


Lambert's cosine law:  $\mathbf{B} = \mathbf{B}_0 \cos(\theta)$ 

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• Number of photons reflected to a given angle  $\theta$  is proportional to  $\cos(\theta)$   $\mathbf{B}_0$ 



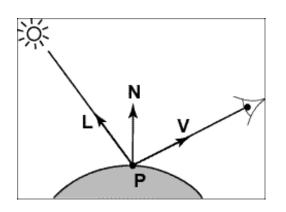
Lambert's cosine law:  $\mathbf{B} = \mathbf{B}_0 \cos(\theta)$ 

 But appearance is the same from every angle due to larger pixel footprint at larger angles

$$\mathbf{A}_{ heta} \propto \mathbf{A}_0 \frac{1}{\cos \theta}$$

Radiance (what eye sees) 
$$\propto \mathbf{B}_0 \mathbf{A}_0 \cos(\theta) \frac{1}{\cos(\theta)}$$

#### Final Lambertian image formation model



$$I = k_d \mathbf{N} \cdot \mathbf{L}$$

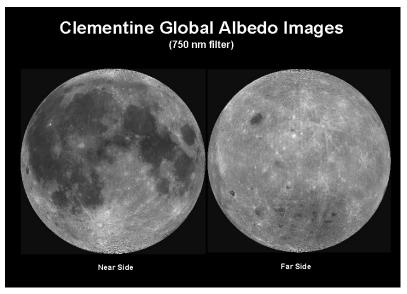


- 1. Diffuse **albedo**: what fraction of incoming light is reflected?
  - Introduce scale factor $k_d$
- 2. Light intensity: how much light is arriving?
  - Compensate with camera exposure (global scale factor)
- 3. Camera response function
  - Assume pixel value is linearly proportional to incoming energy (perform radiometric calibration if not)

#### **Albedo**

#### Sample albedos

Surface	Typical albedo
Fresh asphalt	0.04 <sup>[4]</sup>
Open ocean	0.06 <sup>[5]</sup>
Worn asphalt	0.12 <sup>[4]</sup>
Conifer forest (Summer)	0.08, <sup>[6]</sup> 0.09 to 0.15 <sup>[7]</sup>
Deciduous trees	0.15 to 0.18 <sup>[7]</sup>
Bare soil	0.17 <sup>[8]</sup>
Green grass	0.25 <sup>[8]</sup>
Desert sand	0.40 <sup>[9]</sup>
New concrete	0.55 <sup>[8]</sup>
Ocean ice	0.5–0.7 <sup>[8]</sup>
Fresh snow	0.80-0.90 <sup>[8]</sup>

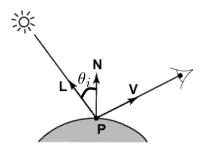


Objects can have varying albedo and albedo varies with wavelength

Source: <a href="https://en.wikipedia.org/wiki/Albedo">https://en.wikipedia.org/wiki/Albedo</a>

### A Single Image: Shape from shading





Suppose (for now)  $k_d=1$ 

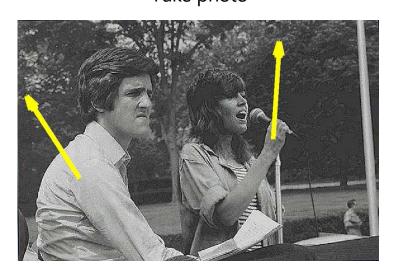
$$I = k_d \mathbf{N} \cdot \mathbf{L}$$
$$= \mathbf{N} \cdot \mathbf{L}$$
$$= \cos \theta_i$$

You can directly measure angle between normal and light source

- Not quite enough information to compute surface shape
- But can be if you add some additional info, for example
  - assume a few of the normals are known (e.g., along silhouette)
  - constraints on neighboring normals—"integrability"
  - smoothness
- Hard to get it to work well in practice
  - plus, how many real objects have constant albedo?
  - But, deep learning can help

## **Application: Detecting composite photos**

Fake photo

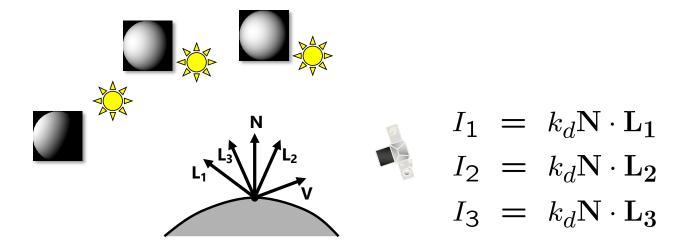


Real photo



## Let's take more than one photo!

#### **Photometric stereo**



Can write this as a matrix equation:

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = k_d \begin{bmatrix} \mathbf{L_1}^T \\ \mathbf{L_2}^T \\ \mathbf{L_3}^T \end{bmatrix} \mathbf{N}$$

#### Solving the equations

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = \begin{bmatrix} \mathbf{L}_1^T \\ \mathbf{L}_2^T \\ \mathbf{L}_3^T \end{bmatrix} k_d \mathbf{N} \qquad \mathbf{G} = \mathbf{L}^{-1} \mathbf{I}$$

$$k_d = \|\mathbf{G}\|$$

$$K_d = \|\mathbf{G}\|$$

$$\mathbf{N} = \frac{1}{k_d} \mathbf{G}$$

$$\mathbf{N} = \frac{1}{k_d} \mathbf{G}$$

Solve one such linear system **per pixel** to solve for that pixel's surface normal

#### More than three lights

Can get better results by using more than 3 lights

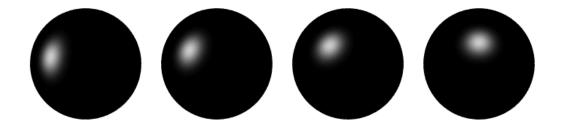
$$\begin{bmatrix} I_1 \\ \vdots \\ I_n \end{bmatrix} = \begin{bmatrix} \mathbf{L_1} \\ \vdots \\ \mathbf{L_n} \end{bmatrix} k_d \mathbf{N}$$

Least squares solution:

$$egin{array}{lll} \mathbf{I} &=& \mathbf{L}\mathbf{G} \\ \mathbf{L}^T\mathbf{I} &=& \mathbf{L}^T\mathbf{L}\mathbf{G} \\ \mathbf{G} &=& (\mathbf{L}^T\mathbf{L})^{-1}(\mathbf{L}^T\mathbf{I}) \end{array}$$
 Solve for N,  $k_d$  as before 
$$&\qquad \qquad \text{What's the size of } \mathbf{L}^T\mathbf{L}?$$

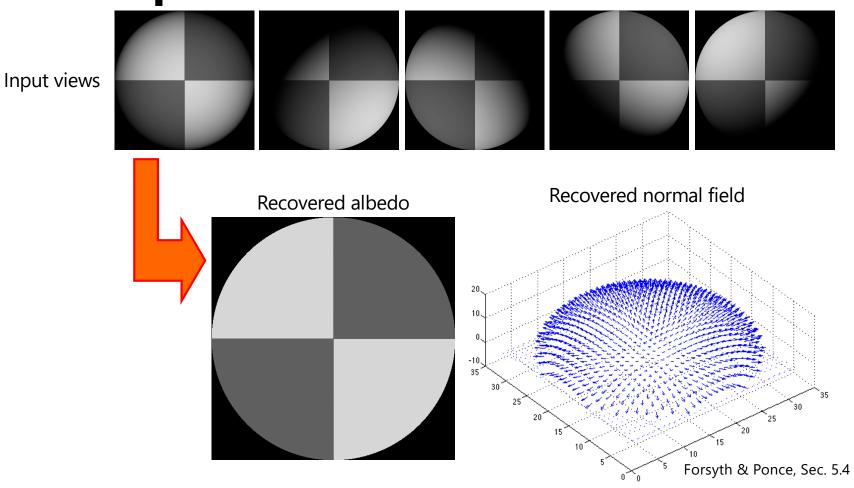
#### Computing light source directions

Trick: place a chrome sphere in the scene



- the location of the highlight tells you where the light source is

**Example** 



### **Depth from normals**

- Solving the linear system perpixel gives us an estimated surface normal for each pixel
- How can we compute depth from normals?
  - Normals are like the "derivative" of the true depth







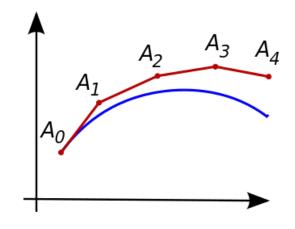
Input photo

Estimated normals

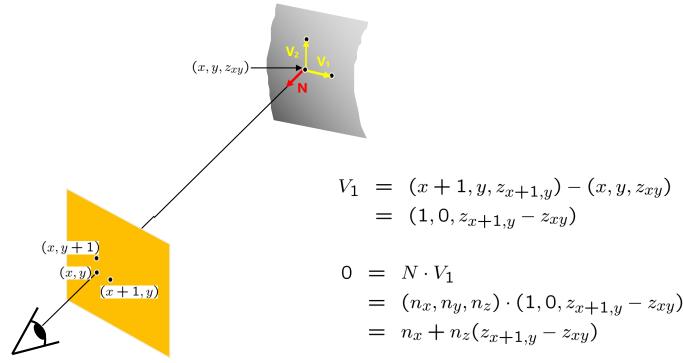
Estimated normals (needle diagram)

#### **Normal Integration**

- Integrating a set of derivatives is easy in 1D
  - (similar to Euler's method from diff. eq. class)
- Could integrate normals in each column / row separately
  - Wouldn't give a good surface
- Instead, we formulate as a linear system and solve for depths that best agree with the surface normals



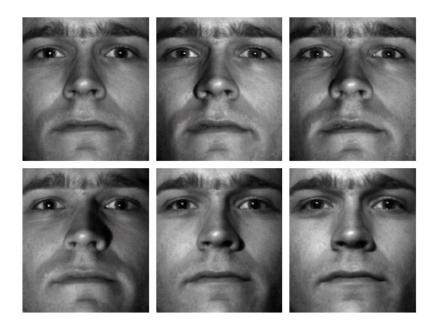
#### **Depth from normals**



#### Get a similar equation for V<sub>2</sub>

- Each normal gives us two linear constraints on z
- compute z values by solving a matrix equation

#### **Results**







### Results

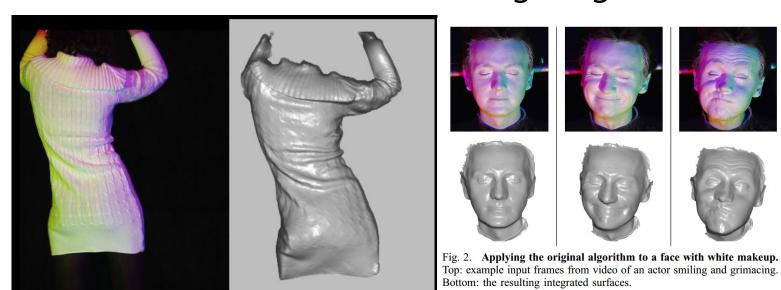






#### **Extension**

Photometric Stereo from Colored Lighting



#### **Video Normals from Colored Lights**

Gabriel J. Brostow, Carlos Hernández, George Vogiatzis, Björn Stenger, Roberto Cipolla <u>IEEE TPAMI</u>, Vol. 33, No. 10, pages 2104-2114, October 2011.

# **Questions?**

# For now, ignore specular reflection





Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

### **And Refraction...**



Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

### And Interreflections...





Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

# **And Subsurface Scattering...**





Slides from Photometric Methods for 3D Modeling, Matsushita, Wilburn, Ben-Ezra

### Limitations

### Bigger problems

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections

### Smaller problems

- camera and lights have to be distant
- calibration requirements
  - measure light source directions, intensities
  - camera response function

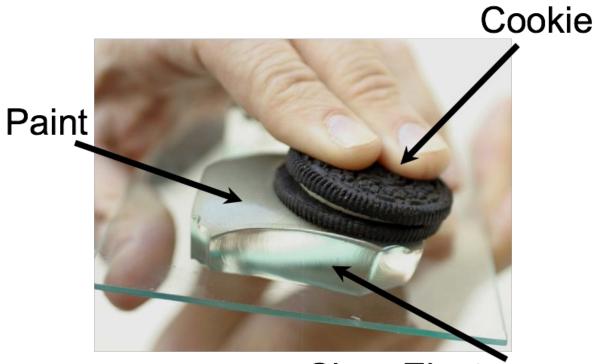
#### Newer work addresses some of these issues

#### Some pointers for further reading:

- Zickler, Belhumeur, and Kriegman, "<u>Helmholtz Stereopsis: Exploiting Reciprocity for Surface Reconstruction</u>." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann & Seitz, "<u>Example-Based Photometric Stereo: Shape Reconstruction with General, Varying BRDFs</u>." IEEE Trans. PAMI 2005



Johnson and Adelson, 2009



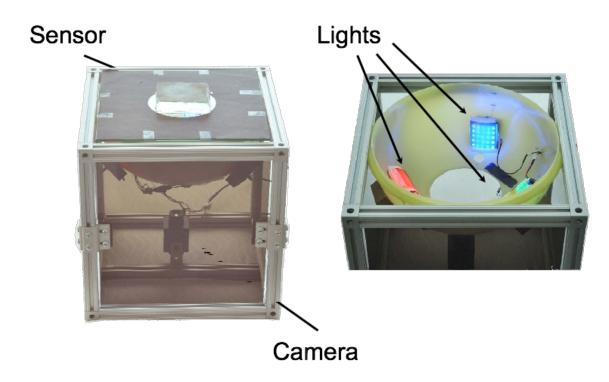
Clear Elastomer

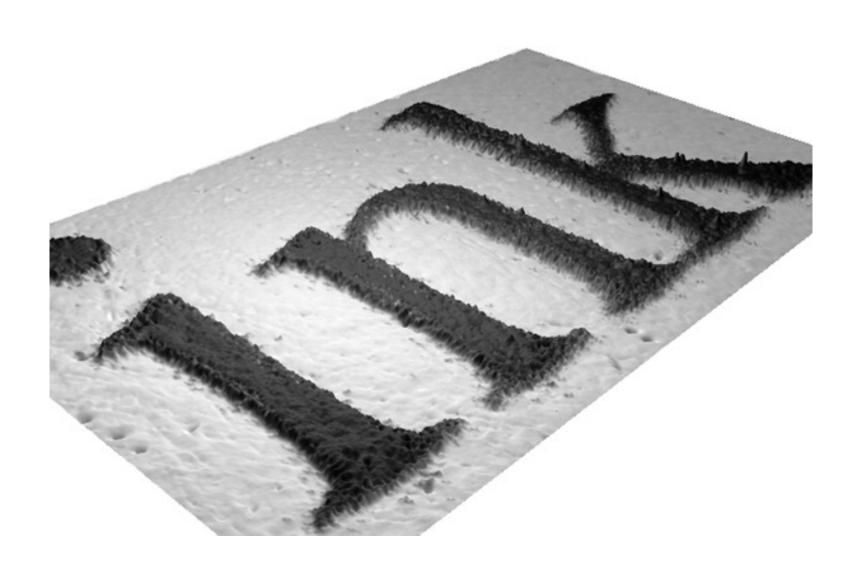
Johnson and Adelson, 2009

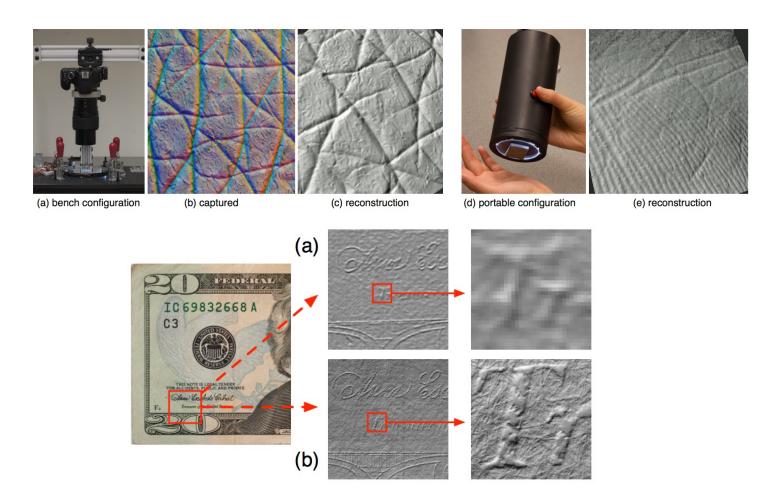




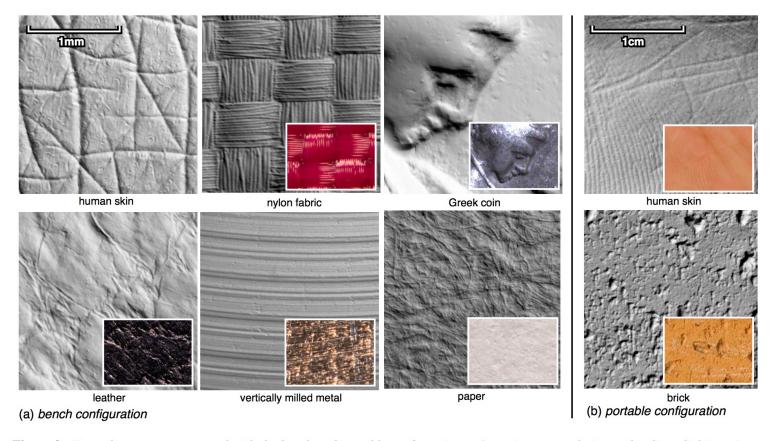
## Lights, camera, action



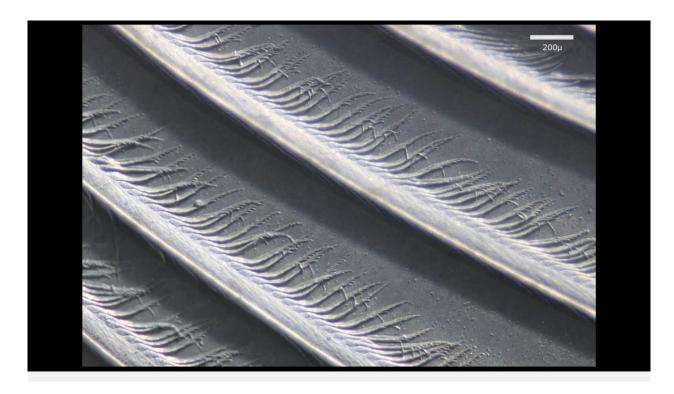




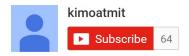
**Figure 7:** Comparison with the high-resolution result from the original retrographic sensor. (a) Rendering of the high-resolution \$20 bill example from the original retrographic sensor with a close-up view. (b) Rendering of the captured geometry using our method.



**Figure 9:** Example geometry measured with the bench and portable configurations. Outer image: rendering under direct lighting. Inset: macro photograph of original sample. Scale shown in upper left. Color images are shown for context and are to similar, but not exact scale.



Sensing Surfaces with GelSight



138,850 views

https://www.youtube.com/watch?v=S7gXih4XS7A

#### **InverseRenderNet:** Learning single image inverse rendering

#### Ye Yu and William A. P. Smith Department of Computer Science, University of York, UK

{yy1571,william.smith}@york.ac.uk

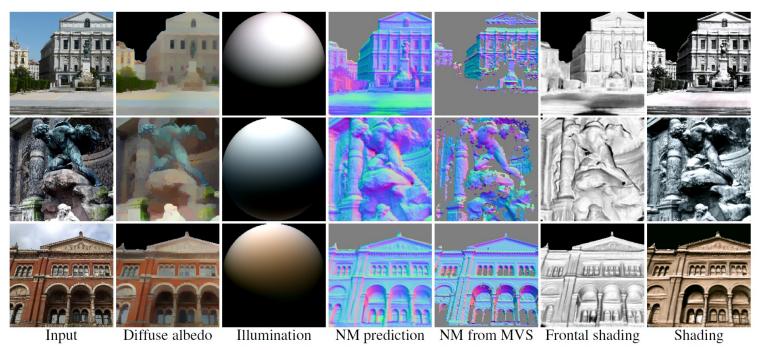


Figure 1: From a single image (col. 1), we estimate albedo and normal maps and illumination (col. 2-4); comparison multiview stereo result from several hundred images (col. 5); re-rendering of our shape with frontal/estimated lighting (col. 6-7).

# **Questions?**